

Greenhouse Gasses/Alternate Energy
next class: in computer lab A119(off
the caf)

Blackbody radiation - total and peak
emission
absorption factors

Blackbody - an ideal object that
absorbs all wavelengths and emits all
wavelengths of electromagnetic
radiation.

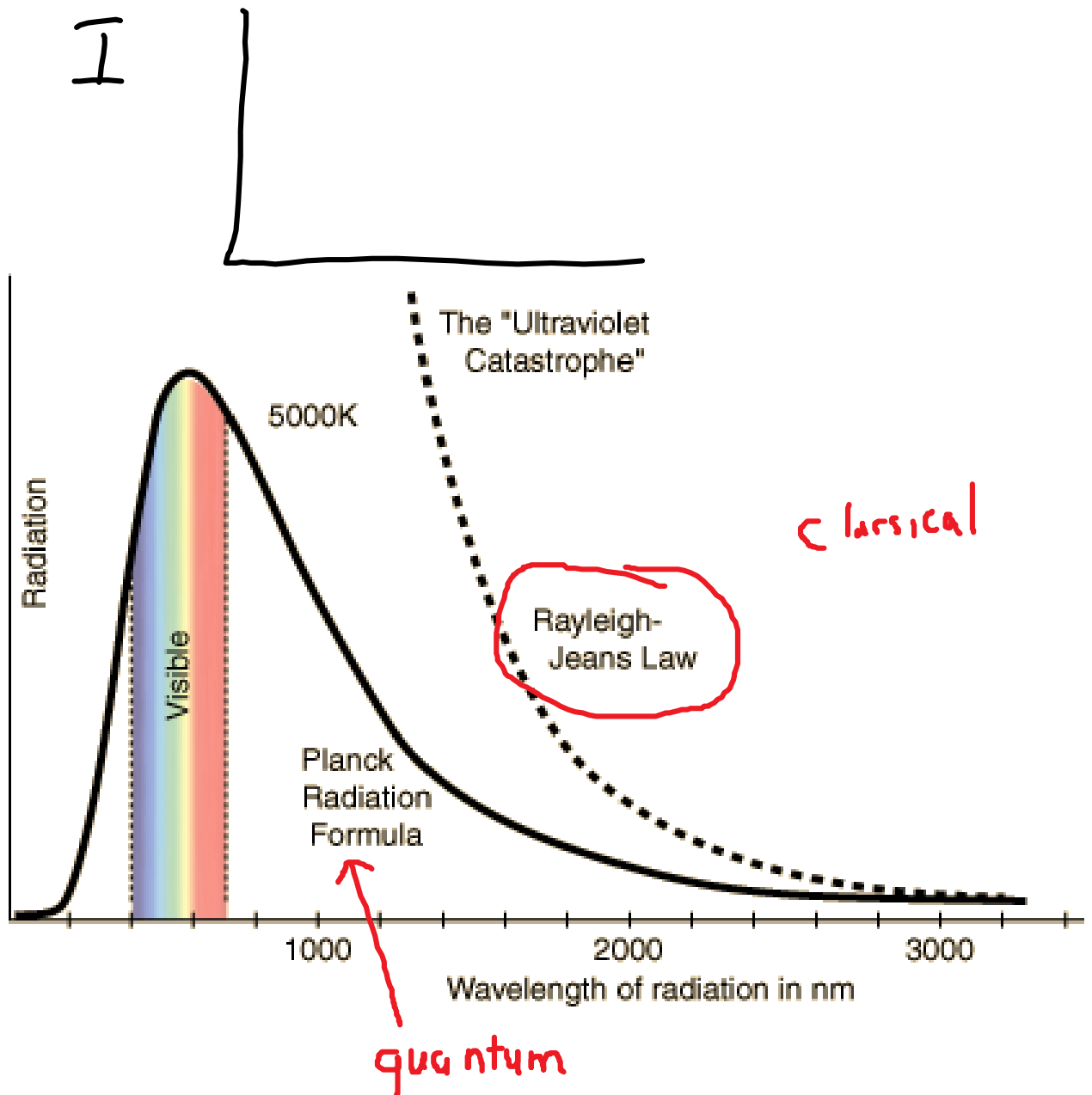
late 1800s Max Planck was hired by
the German government to build a
better lightbulb

used Maxwell's equations of
electromagnetism and statistical
mechanics to derive a relationship.

It didn't match observation.

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it didn't match observation.



Planck assumed that EM radiation is quantized by $E=hf$ energy of a photon
 $h = 6.63 \times 10^{-34} \text{Js}$
the new graph matched observation

The area under the graph corresponds to the total power output of the blackbody

$$P = \varepsilon \sigma A T^4$$

P is the total power output, in Watts.

A is the surface area of the object, in m^2 . Sphere $A=4\pi r^2$

T is the temperature of the blackbody.
must be in Kelvin!!!! +273 to get Kelvin from Celsius

σ is Stefan-Boltzmann constant = $5.67 \times 10^{-8} \text{ Wm}^{-2}\text{K}^{-4}$

ε is emissivity for the object to be not a perfect blackbody. No units.

- emission coefficient -

the peak emission is dependent on the temperature.

$$\lambda_{\text{max}} = (2.90 \times 10^{-3} \text{ mK})/T$$

I, intensity = power/Area

albedo = total scattered power/(total incident power)

factors that influence albedo are:
colour, reflectivity (amount of ice/water
vs earth), texture,

eg.

The Sun has a radius of 695500km
and a surface temperature of 5000K.

a) what is the total power output of the
sun if it is an ideal blackbody $\varepsilon=1$

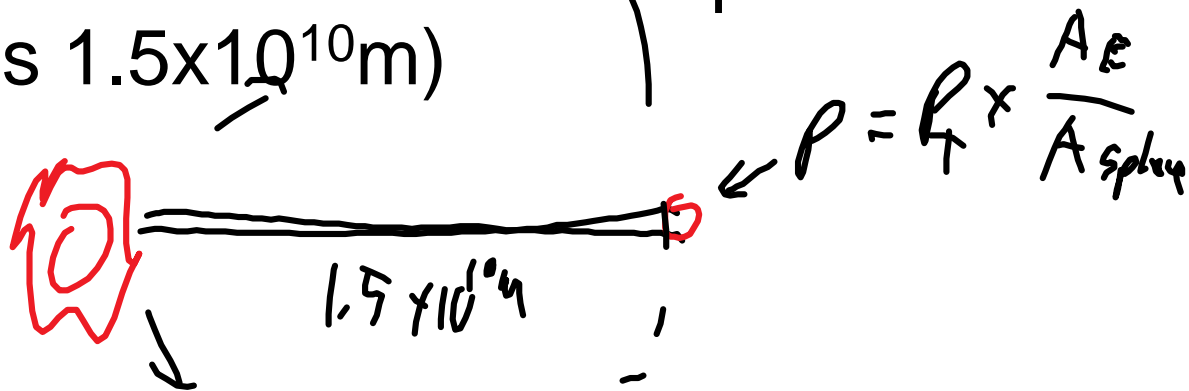
$$P = \varepsilon \sigma AT^4$$

$$P = 1 \times 5.67 \times 10^{-8} \text{ Wm}^{-2}\text{K}^{-4} \\ \times (4\pi \times (6.955 \times 10^8 \text{ m})^2) \times (5000 \text{ K})^4 \\ 2.15 \times 10^{26} \text{ W}$$

b) The Earth has a radius of $6.371 \times 10^6 \text{ m}$ and an average surface temperature of 15°C and an emissivity of 0.5. What is the total power output of the Earth?

$$\begin{aligned}
 P &= 0.5 \times 5.67 \times 10^{-8} \text{ Wm}^{-2}\text{K}^{-4} \\
 &\times (4 \times \pi \times (6.371 \times 10^6 \text{ m})^2) \times (15 + 273 \text{ K})^4 \\
 &= 9.95 \times 10^{16} \text{ W} \\
 0.5 \times 5.67 \text{ E-}8 \times 4 \times 3.14159 \times 6.37 \text{ E}6^2 \\
 \times 288^4 &= 9.94515 \text{ E}16
 \end{aligned}$$

a) The Sun is $1.5 \times 10^{10} \text{ m}$ away from the Earth. Calculate the energy per unit area (intensity) of sunlight on Earth assuming solution from a. (think of the Earth as a disc on a sphere radius $1.5 \times 10^{10} \text{ m}$)



$$A_E = \text{disc} = \pi r^2 = 3.14159 \times (6.371 \text{ E}6^2) = 1.27516 \text{ E}14$$

$$A_{\text{sphere}} = 4\pi r^2 = 4 \times 3.14159 \times (1.5 \text{ E}10^2) = 2.82743 \text{ E}21$$

$$P =$$

$$2.15 \text{ E}26 \times 1.27516 \text{ E}14 / 2.82743 \text{ E}21 = 9.69642 \text{ E}18 \text{ W power into the earth}$$

if that spread over the disc
 $9.69642\text{E}18/1.27516\text{E}14=$
 $76,040.81056 = 76\text{kW/m}^2$

- a) What albedo would make the solution to b (output energy) match the solution to c (input energy)?
- b) what is the peak emission wavelength for Sun and the Earth?